

Lecture 12

(Poynting Vector and Standing Waves)

Physics 2310-01 Spring 2020

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Review of Power in Waves

- Let's examine a wave moving to the right on a string.
- At point a, the force from the string to the left, F , can be broken down into the horizontal and vertical components using the slope of the string:

$$F_y(x, t) = -F \frac{\partial y(x, t)}{\partial x}$$

- And then the power at a is just given by the force times the velocity of the piece of string at that point:

$$P(x, t) = F_y(x, t) v_y(x, t) = -F \frac{\partial y(x, t)}{\partial x} \frac{\partial y(x, t)}{\partial t}$$

- Since, $y(x, t) = A \cos(kx - \omega t)$,
- Then,

$$\frac{\partial y(x, t)}{\partial x} = -kA \sin(kx - \omega t) \quad \frac{\partial y(x, t)}{\partial t} = \omega A \sin(kx - \omega t)$$

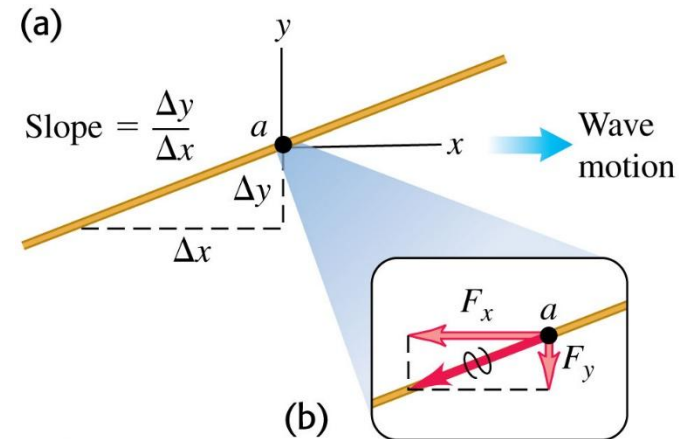
- And the power is given by:

$$P(x, t) = -Fk\omega A^2 \sin^2(kx - \omega t)$$

- Thus, the time averaged power is given by:

$$P_{\text{Avg}} = \frac{1}{2} Fk\omega A^2$$

- Notice that it depends on the amplitude squared



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Power in an EM Wave

- Let's review what we know about the energy stored in fields:

$$u_E = \frac{1}{2} \epsilon_0 E^2, \quad u_B = \frac{1}{2} \frac{1}{\mu_0} B^2$$

- So, at any point in the electromagnetic wave,

$$u = u_E + u_B = \frac{1}{2} \epsilon_0 E^2 + \frac{1}{2} \frac{1}{\mu_0} B^2$$

- But, $E = cB \Rightarrow B = \frac{E}{c}$ 

- So
$$u = \frac{1}{2} \epsilon_0 E^2 + \frac{1}{2} \frac{1}{\mu_0 c^2} E^2 = \frac{1}{2} \epsilon_0 E^2 + \frac{1}{2} \epsilon_0 E^2 = \epsilon_0 E^2$$

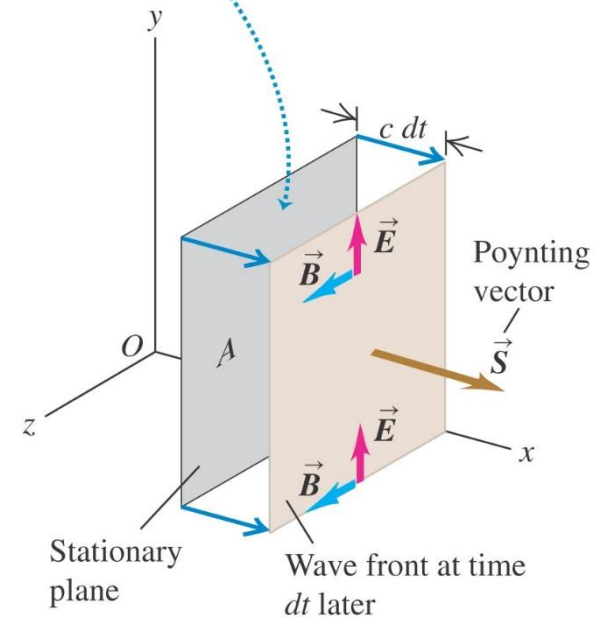
$$\mu_0 c^2 = \frac{\mu_0}{\mu_0 \epsilon_0} = \frac{1}{\epsilon_0}$$

Power in an EM Wave

- But this is the energy density at a localized point in space. We want to describe how this energy moves with the wave.
- Let's examine the energy (dU) in a plane wave that passes through a certain area (A) within a certain time (dt).

$$dU = u dV = uA(c dt) \Rightarrow$$
$$dU = (\epsilon_0 E^2) A c dt$$

At time dt , the volume between the stationary plane and the wave front contains an amount of electromagnetic energy $dU = uAc dt$.



Poynting Vector

$$dU = (\epsilon_0 E^2) A c dt$$

- Now, let us define the **energy per unit time, per unit area** as:

$$S = \frac{1}{A} \frac{dU}{dt} = c \epsilon_0 E^2$$

- Notice that it scales with the square of the field (wave amplitude).
- Now, just for fun, let's rewrite this using $E = cB$:

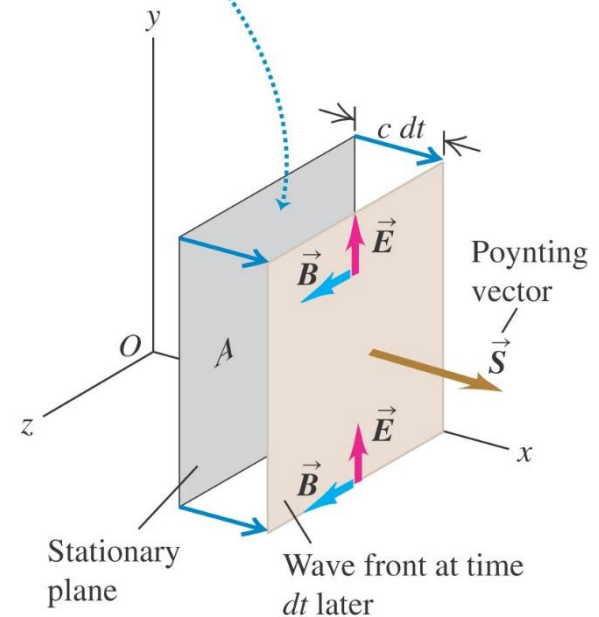
$$S = c^2 \epsilon_0 EB = \frac{1}{\mu_0} EB$$

- And put it in vector form to denote the direction of energy transmission:

$$\vec{S} = \frac{1}{\mu_0} \vec{E} \times \vec{B}$$

Poynting Vector

At time dt , the volume between the stationary plane and the wave front contains an amount of electromagnetic energy $dU = uAc dt$.



Poynting Vector

- Remember, the Poynting vector represents the ***energy per time per unit area*** instantaneously passing through an area at a particular point in space.

$$S = \frac{\text{energy}}{\text{time} \cdot \text{area}} = \frac{\text{power}}{\text{area}}$$

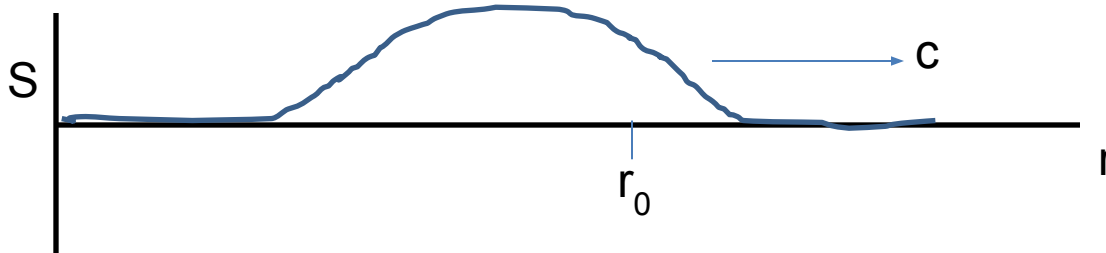
- So, if we want to know the entire *instantaneous* power output from a source, we have to integrate the Poynting vector over a closed surface containing the source:

$$P = \oint \vec{S} \cdot d\vec{A}$$

- We'll return to this in a moment, but let's first get rid of that annoying term: instantaneous.

Average Power (per unit area)

- The time average of the Poynting vector will depend on its time dependence, and over what length of time you average.
- For instance, look at the following pulse:



- If we integrate over the time from when the pulse just reaches r_0 until the maximum reaches r_0 , we will get one answer.
- If we integrate over all time, we will get an answer that will approach zero...
- So, when we speak of average power, we either have to be precise in our description of the time integration, or...

Average Power (per unit area)

- Assume a never-ending sinusoidal plane wave.

$$\begin{aligned}\vec{S} &= \frac{1}{\mu_0} \vec{E}(x,t) \times \vec{B}(x,t) \\ &= \frac{1}{\mu_0} \left[\hat{j} E_{\text{Max}} \sin(kx - \omega t) \right] \times \left[\hat{k} B_{\text{Max}} \sin(kx - \omega t) \right] \\ &= \frac{\hat{i}}{\mu_0} E_{\text{Max}} B_{\text{Max}} \sin^2(kx - \omega t)\end{aligned}$$

- Then, because the average of \sin^2 is $\frac{1}{2}$,

$$I = \left| \vec{S}_{\text{Avg}} \right| = \frac{1}{2\mu_0} E_{\text{Max}} B_{\text{Max}} = \frac{E_{\text{Max}}^2}{2\mu_0 c} = \frac{1}{2} \epsilon_0 c E_{\text{Max}}^2$$

- Where, I is the intensity of the EM wave.

Example

- Consider a source of EM waves, say a light bulb that puts out 100W spread equally in all directions. What is the value of the intensity of light 10m from the bulb?

$$\frac{P}{A(r)} = \frac{100W}{4\pi(10m)^2} = \frac{1W}{4\pi m^2} = S_{\text{Avg}} = I$$

- But remember that the intensity is proportional to the amplitude squared:

$$I = \left| \vec{S}_{\text{Avg}} \right| = \frac{1}{2\mu_0} E_{\text{Max}} B_{\text{Max}} = \frac{E_{\text{Max}}^2}{2\mu_0 c} = \frac{1}{2} \epsilon_0 c E_{\text{Max}}^2$$

- So, as pointed out earlier, for spherical waves, $\left| \vec{E}(r, t) \right| = \frac{E_0}{r} \sin(kr - \omega t)$

Electromagnetic Momentum Flow

- Electromagnetic waves also carry momentum^{*}, with a momentum density:

$$\frac{dp}{dV} = \frac{EB}{\mu_0 c^2} = \frac{S}{c^2}$$

- And you can then consider a momentum flow rate (momentum passing through unit area in a period of time dt):

$$\frac{1}{A} \frac{dp}{dt} = \frac{EB}{\mu_0 c} = \frac{S}{c}$$

- Since $dV = Acdt$

^{*}see <http://farside.ph.utexas.edu/teaching/em/lectures/node90.html> for a discussion of how we know this.

Radiation Pressure

- And the average momentum flow is just:

$$\left\langle \frac{1}{A} \frac{dp}{dt} \right\rangle = \frac{S_{\text{Avg}}}{c} = \frac{I}{c}$$

- Now, remember Newton's second law:

$$\vec{F} = \frac{d\vec{p}}{dt}$$

- And since pressure is just force per area,

$$\left\langle \frac{1}{A} \frac{dp}{dt} \right\rangle = \left\langle \frac{F}{A} \right\rangle = \langle p_{\text{rad}} \rangle = \frac{S_{\text{Avg}}}{c} = \frac{I}{c}$$

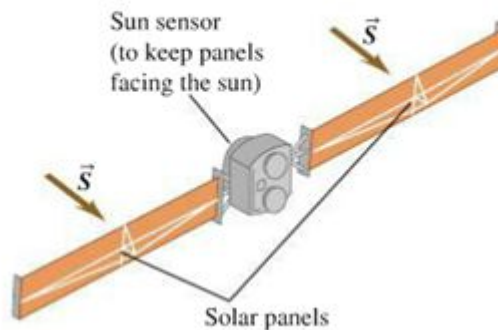
- If the wave is reflected, it's twice this.

Example

Example 32.5 Power and pressure from sunlight

An earth-orbiting satellite has solar energy–collecting panels with a total area of 4.0 m^2 (Fig. 32.21). If the sun’s radiation is perpendicular to the panels and is completely absorbed, find the average solar power absorbed and the average radiation-pressure force.

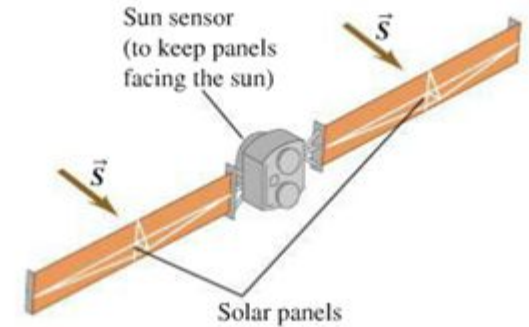
32.21 Solar panels on a satellite.



The intensity I (power per unit area) is $1.4 \times 10^3 \text{ W/m}^2$.

Example

32.21 Solar panels on a satellite.

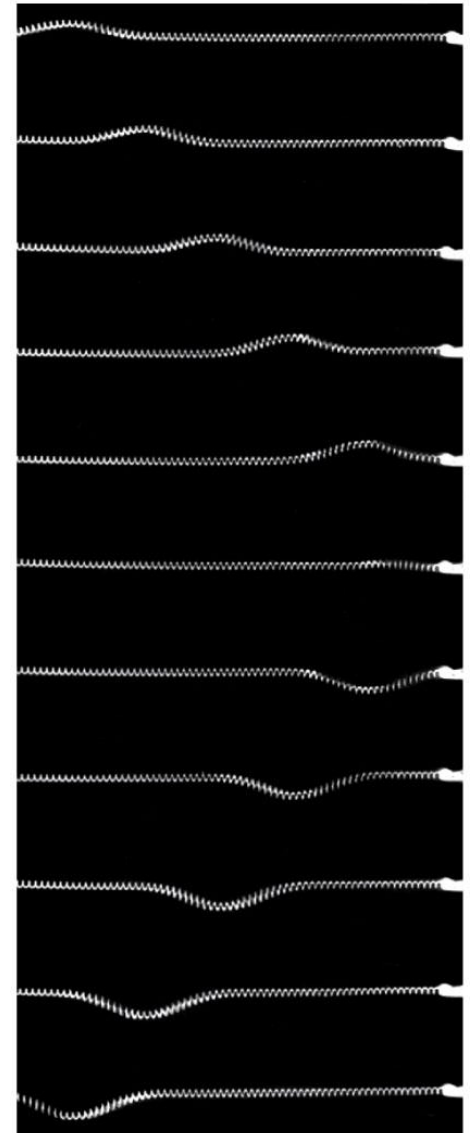


The intensity I (power per unit area) is $1.4 \times 10^3 \text{ W/m}^2$.

$$p_{\text{rad}} = \frac{I}{c} = \frac{1.4 \times 10^3 \text{ W/m}^2}{3.0 \times 10^8 \text{ m/s}} = 4.7 \times 10^{-6} \text{ Pa}$$

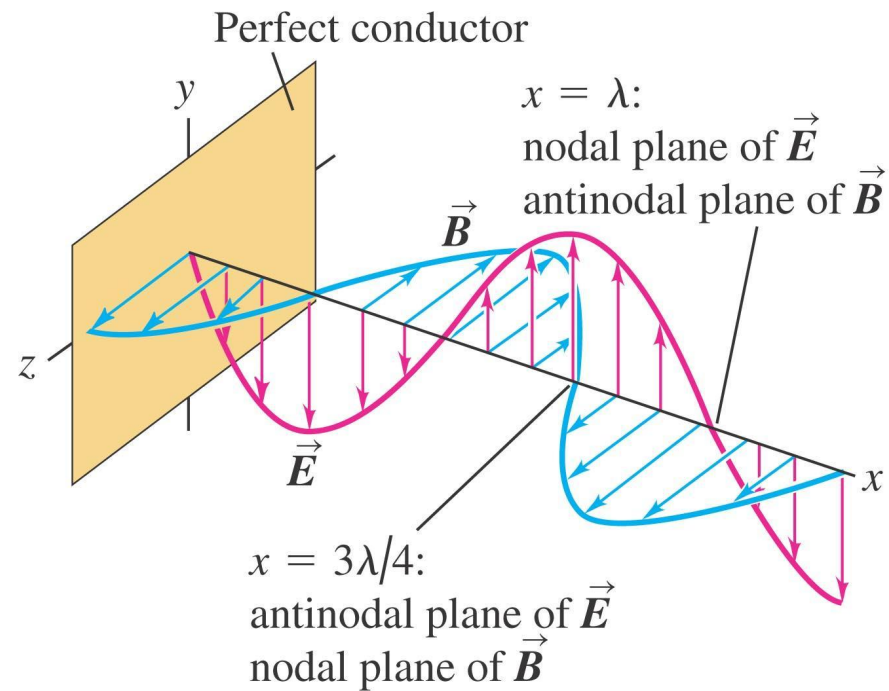
Reflections at Conducting Surfaces

- What happens when a transverse matter wave strikes a fixed boundary?
- The wave is reflected and inverted.
- A similar thing happens when an EM wave strikes a conducting boundary...



Reflections at Conducting Surfaces

- The electric field of the wave causes charges to move on the surface of the conductor.
- The net effect of their motion is to keep the electric field inside the conductor **zero**.
- The moving charges also cause a magnetic field.
- Hence, these charges create an electromagnetic wave which is inverted from the original and travels in the opposite direction, away from the conductor.



Standing EM Waves

- Take the incoming wave (moving in the $-x$ direction) as:

$$\vec{E}_{in}(x,t) = \hat{j}E_{Max} \cos(kx + \omega t), \quad \vec{B}_{in}(x,t) = -\hat{k}B_{Max} \cos(kx + \omega t)$$

- Then the reflected wave must be given by:

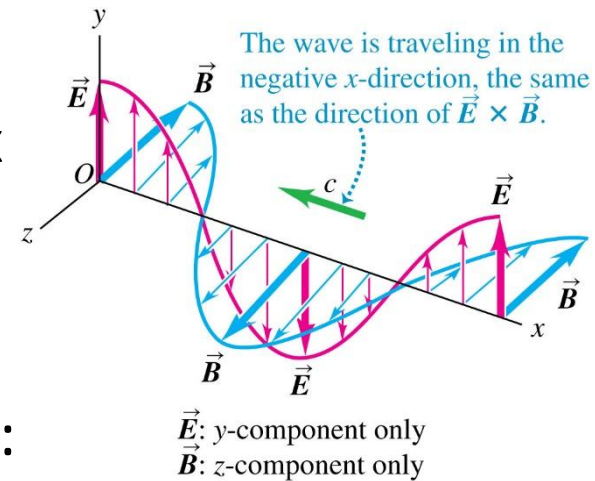
$$\vec{E}_{ref}(x,t) = -\hat{j}E_{Max} \cos(kx - \omega t), \quad \vec{B}_{ref}(x,t) = -\hat{k}B_{Max} \cos(kx - \omega t)$$

- The resulting superposition of these two is:

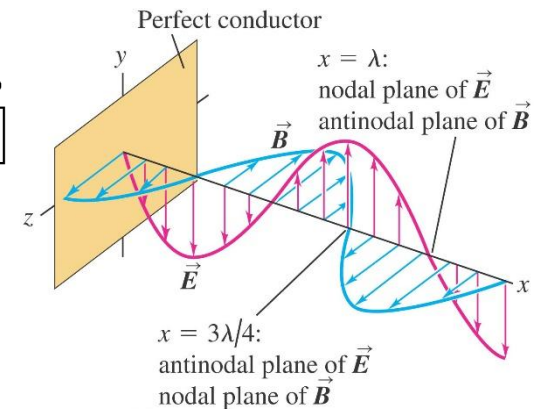
$$\begin{aligned} \vec{E}_{tot}(x,t) &= \vec{E}_{in}(x,t) + \vec{E}_{ref}(x,t) = \hat{j}E_{Max} [\cos(kx + \omega t) - \cos(kx - \omega t)], \\ B_{tot}(x,t) &= B_{in}(x,t) + B_{ref}(x,t) = -\hat{k}B_{Max} [\cos(kx + \omega t) + \cos(kx - \omega t)] \end{aligned}$$

- Or, simplifying:

$$\begin{aligned} \vec{E}_{tot}(x,t) &= -\hat{j}2E_{Max} \sin(kx) \sin(\omega t), \\ B_{tot}(x,t) &= -\hat{k}2B_{Max} \cos(kx) \cos(\omega t) \end{aligned}$$



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Standing EM Waves

- There are then nodes in both the electric field at:

$$x = 0, \frac{\lambda}{2}, \lambda, \frac{3\lambda}{2} \dots$$

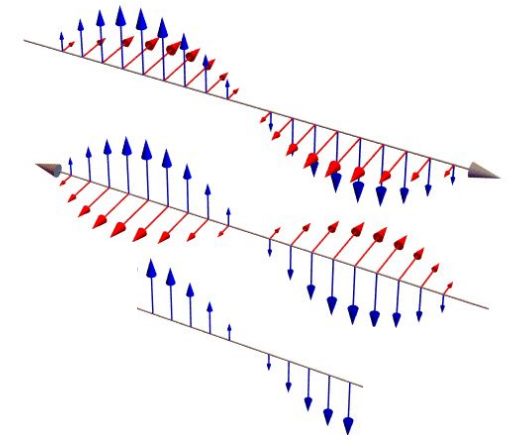
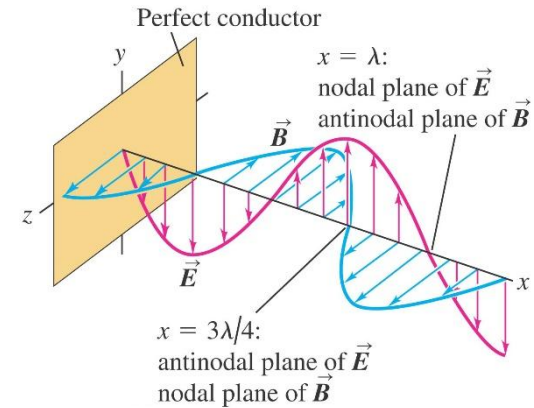
- And in the B-field at:

$$x = \frac{\lambda}{4}, \frac{3\lambda}{4}, \frac{5\lambda}{4} \dots$$

- Note also that the electric and magnetic field are out of phase *in time*

$$\begin{aligned} \vec{E}_{\text{tot}}(x, t) &= -\hat{j}2E_{\text{Max}} \sin(kx) \sin(\omega t), \\ B_{\text{tot}}(x, t) &= -\hat{k}2B_{\text{Max}} \cos(kx) \cos(\omega t) \end{aligned}$$

- When ωt is such that the electric field is zero everywhere, the magnetic field is maximum everywhere!



B is blue, E is red

Standing EM Waves

- What if we put in another conducting plane, we create an electromagnetic cavity that will sustain a resonance for EM waves of wavelength:

$$\lambda_n = \frac{2L}{n}, \quad (n = 1, 2, 3 \dots)$$

