Physics 480/581

Problem Session No. 5

Monday, 1 October, 2018

1. Having determined all the equations for a geodesic in the Schwarzschild metric, using our usual orthonormal tetrad, you will know that the last of those 4 equations was the following:

$$\frac{d}{d\tau} \left(\mathcal{H} \frac{dt}{d\tau} \right) + \mathcal{H}' \left(\frac{1}{\mathcal{H}} \frac{dr}{d\tau} \right) \left(\mathcal{H} \frac{dt}{d\tau} \right) = 0 \ .$$

Please show that this can be re-written as a "perfect derivative," i.e., simply the action of $d/d\tau$ on some expression. Therefore you will have discovered a *constant of the motion* for masses moving in this manifold.

2. Consider the second component of the covariant derivative, in the $\hat{\theta}$ direction:

$$\frac{d}{d\tau} \left(r \frac{d\theta}{d\tau} \right) + \frac{\mathcal{H}}{r} \left(\frac{1}{\mathcal{H}} \frac{dr}{d\tau} \right) r \frac{d\theta}{d\tau} - \frac{\cot \theta}{r} \left(r \sin \theta \frac{d\varphi}{d\tau} \right)^2 = 0.$$

It's a second-order differential equation, so one may give it two initial conditions. Show that if one sets $\theta = \pi/2$, the equatorial plane, and also $d\theta/d\tau = 0$ at the initial time, then it will always stay in the equatorial plane.

3. Use the given connections to determine the following two curvature 2-forms, $\Omega_{\hat{r}\hat{\theta}}$ and $\Omega_{\hat{\theta}\hat{\varphi}}$, via the generic expression:

$$\Omega_{\lambda\mu} = d\Gamma_{\lambda\mu} + \Gamma_{\lambda\nu} \wedge \Gamma^{\nu}_{\mu} .$$

4. Using the Schwarzschild metric, imagine a circle in the equatorial plane, with a constant value of r > 2M and $\theta = \pi/2$, at a fixed time. What is the radius of this circle?