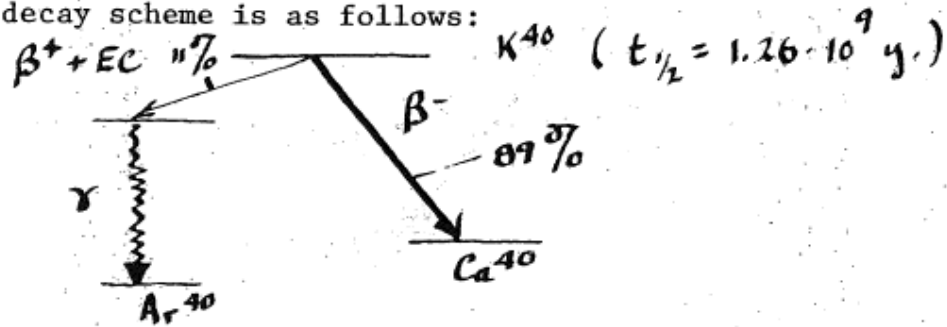


10 pts

- (1) Consider the problem of determining the age of a sample of the earth (a rock) from the relative abundance of  $K^{40}$  and  $Ar^{40}$  in that rock. We may safely assume that while the material is molten or gaseous, the  $Ar^{40}$  escapes. Thus  $Ar^{40}$  slowly builds up from zero abundance starting at the time the rock material last solidified. The relevant decay scheme is as follows:



If one studies a large sample of rocks, one finds that the molar abundance ratio  $\frac{A}{K}$  of  $Ar^{40}$  to  $K^{40}$  is always less than a limiting value:

$$\left( \frac{A}{K} \right)_{\max} = 1.23$$

Using this fact and the decay scheme above, first derive a general formula which expresses the age of the rock  $t$  in terms of  $\frac{A}{K}$  and the partial decay rates to  $Ar^{40}$ ,  $Ca^{40}$ ; then compute the maximum rock age  $t_{\max}$  ( $= t_{\text{EARTH}}$ ).

(2) In 1987 a Supernova went off in the Large Magellanic Cloud about 50 kpc away from us. What has made SN 1987a (the "name" of the supernova) most remarkable is that a burst of neutrinos ( $\nu$ ) were detected by two deep underground detectors, the IMB (in the U.S.) and Kamiokande (in Japan) detectors. 8 events were detected in IMB and 11 at Kamiokande. The properties of these bursts were:

- Spread in  $\nu$  energies observed:

$$\delta E_\nu \sim 10 \text{ MeV}$$

- Average energy:

$$\langle E_\nu \rangle \sim (2-3) \delta E_\nu \sim 20-30 \text{ MeV}$$

- Spread in arrival time:

$$\delta t \sim 10 \text{ s} \quad (\text{is it upper or lower?})$$

Using this data, place a bound on the  $\nu$  mass.

Hints: You do not need to know about oscillations. Start with special relativistic relations between mass, energy and momentum. This is an exercise in approximations!