## ASTRO 536 HW #4 - Due 11/8/17 in class

## Problems/exercises for the Virial Theorem, Hydrostatic Equil.: Problem 1 is for 20 points, half for extra credit; The rest are worth 10 points each

1) Start with the Equation of Motion:
PDV = - PP-PPØ + LJXB
for a cloud of gas with density p, velocity i,
Pressure P, grav. potential Ø, electric current density
J, and magnetic field B. Using Ampere's Caw
without displacement currents, Show that:
$9\frac{DU}{Dt} = -\nabla P - 9\nabla \beta - 1 \nabla (B^2) + 1 (B \cdot \nabla) B$
From this equation derive the Virial Theorem by
taking the scalar product of both sides with it and
integrating over a volume V bounded by surface S.
You should get:
$\frac{1}{2} \frac{d^{2}I}{dt^{2}} = 2J + 3U + m + \Omega + 6(\vec{r} \cdot \vec{B})\vec{B} \cdot d\vec{s}$
Where - (P+B <sup>2</sup> )r.ds /
$I = \left( g r^2 dV \right) $
$U = \int P dV$ $M = I \int B^2 dV$
T-1(01+2/V)
$\Omega = - \left( \frac{1}{2} \nabla \nabla$

2) Consider a gaseous spheu of uniform temperature

T and density g, radius R, in a medium of

constant pressure P. From the virial theorem,

show that such a sphere can be in equilibrium

with the external medium only if the pressure

is less than a critical value:  $P = \frac{3^4 \cdot 10^3}{2^{13}} \left( \frac{1}{10^4} \frac{1}{10^4}$ 

(3) (1) Consider the equation of hydrostatic:

equilibring for a star of radius R,

mass M.

(a) Show that the pressure at the center, P(o),

(50+5) is quien by the equation:

P(o) = 1 (411) for (9(x)) 3 M 3(x) a M(x).

Here, P(x) is the average density of a

sphere of radius r < R inside the star.

(b) Assume that P(x) is a monotonically

(5 ets) decreasing function of r. Prove that:

2 (6 M) \( \text{P}(o) \leq 1 \) Gr(\(\frac{4\pi}{3}\) \( \frac{5}{3}\) \(\frac{6\pi}{3}\) M \(\frac{7}{3}\) (0) M \(\frac{7}{3}\)

En (P(o) \(\frac{1}{3}\) \(\frac{7}{3}\) \(\frac{7}{3}\)