

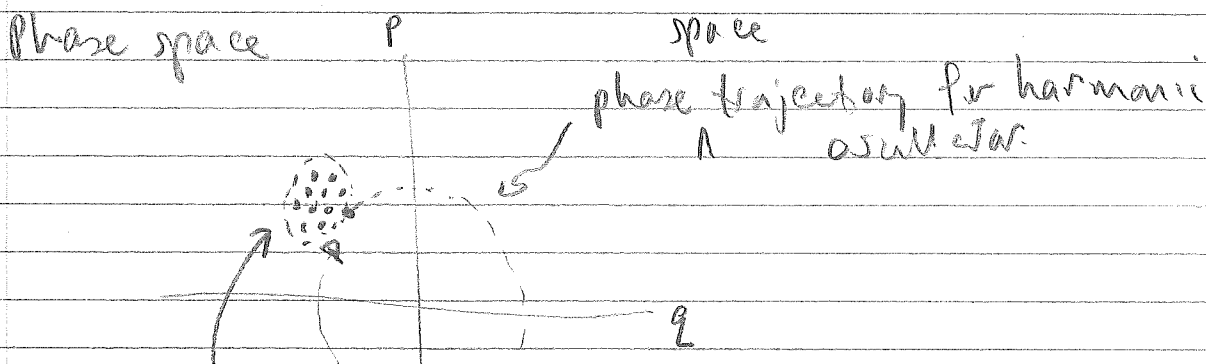
①

(Notes from Taylor)

### Liouville's phase space theorem.

Recall: the canonical equations  $\dot{q}_i = \frac{\partial H}{\partial p_i}$   $\dot{p}_i = -\frac{\partial H}{\partial q_i}$

Where  $H = \sum p_i \dot{q}_i - \mathcal{L}$ , usually  $H = \text{energy}$ .



Consider lots of oscillators starting with nearby initial conditions, occupying area  $A$  in phase space.

How much area will they occupy later?

We can think of phase space filled with a moving fluid. Each "molecule", at  $t=0$ , is an initial condition,  $p_0, q_0$ .

The change in a small volume of fluid (flowing) is

$$dV = \left( \int_S \vec{v} \cdot \hat{n} \, d\vec{a} \right)$$

$\vec{v}$  velocity  $\perp$  to surface

$$\frac{dV}{dt} = \int_S \vec{v} \cdot \hat{n} \, d\vec{a} = \int_V \vec{\nabla} \cdot \vec{v} \, dV \quad \text{Divergence thm.}$$

In phase space, " $\vec{v}$ " =  $(\dot{q}, \dot{p})$  so  $\vec{\nabla} \cdot \vec{v} = \frac{\partial \dot{q}}{\partial q} + \frac{\partial \dot{p}}{\partial p}$

$$= \frac{\partial}{\partial t} \left( \frac{\partial H}{\partial p_i} \right) - \frac{\partial}{\partial p} \left( \frac{\partial H}{\partial q_i} \right) = 0.$$

(2)

7-35 Gravity.

$$p = mv_z$$

$$\frac{dp}{dt} = ma_z = -mg$$

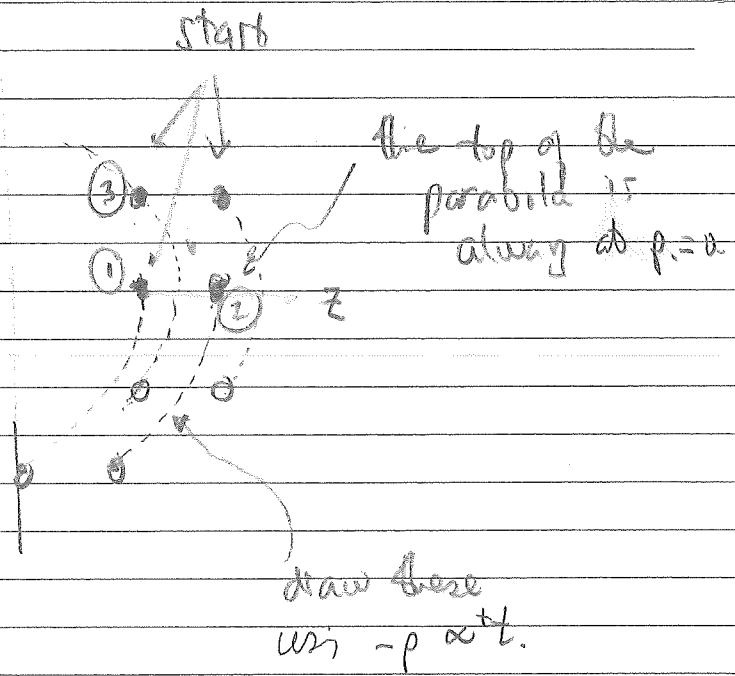
$$p = p_0 - mgt$$

$$z = z_0 + v_0 t - \frac{1}{2}gt^2$$

Now  $v_0 = \frac{p_0}{m}$

$v_0 = 0, p \propto t.$

p



Area starts as rectangle, becomes parallelogram

It will be

$$\Delta z(t) \times \Delta p(t)$$

$\Delta z:$

$$z_1 = z_0 - \frac{1}{2}gt^2$$

$$z_2 = z_0 + \Delta z_0 - \frac{1}{2}gt^2$$

$$\Delta z = \Delta z_0 \text{ for all } t.$$

$\Delta p:$

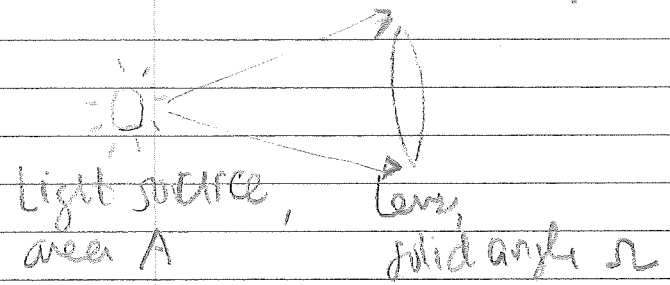
$$p_1 = 0 - mgt$$

(I took  $p_0 = 0$ )

$$p_2 = \Delta p_0 - mgt$$

$$\Delta p = \Delta p_0 \text{ v.t.}$$

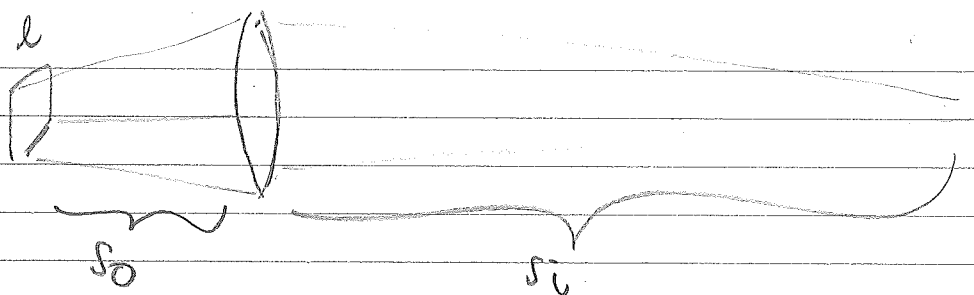
Liouville's theorem is important in stat mech & optics.



angle  $\propto \vec{p}_{\text{transverse}}$

$$\text{So } A\Omega = \text{constant.}$$

(3)



You know  $M = \frac{s_i}{s_o} \approx \frac{l}{s_o}$  so  $A_{\text{obj}} \approx M^2 l^2$

Solid angle subtended by the lens =  $\frac{A_{\text{lens}}}{4\pi s_o^2}$

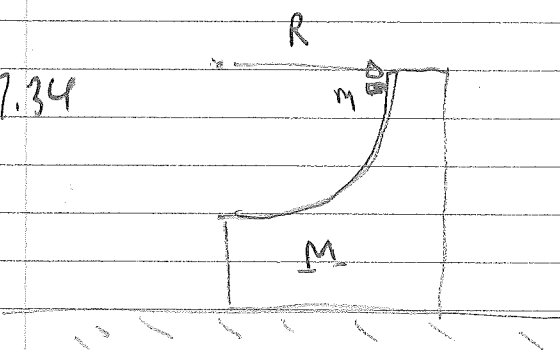
$$\bar{V}_{\phi_o} = \text{Area obj.} \cdot \Omega_o = l^2 \cdot \frac{A_{\text{lens}}}{4\pi s_o^2}$$

$$\bar{V}_{\phi_i} = \text{Area image} \cdot \Omega_i = M^2 l^2 \cdot \frac{A_{\text{lens}}}{4\pi s_i^2} \quad s_i^2 = M^2 s_o^2 \quad \text{so} \quad \bar{V}_{\phi_i} = \bar{V}_{\phi_o}$$

For your microscope, you can get a 50 W lamp with an arc size of 1 mm, or a 200 W lamp with an arc size of 2 mm x 4 mm x 1 mm.

Which is "brighter"?

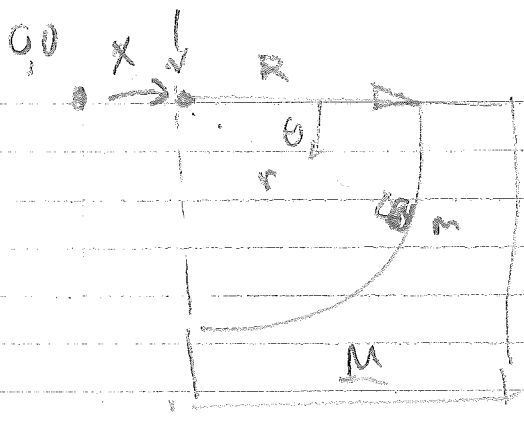
7.34



(1) You don't need Lagrangian mechanics for this. You know that the loss potential energy to m drops  $\rightarrow$  kinetic energy of M and m.

$$\text{Also, } \vec{P}_M + \vec{P}_m = 0$$

Let's use Lagrangian formalism, though.



Wedge of X.

$$\text{mass } m, \quad x_m = x + r \cos \theta$$

$$y_m = -r \sin \theta$$

$$T_m = \frac{1}{2} m (\dot{x}_m^2 + \dot{y}_m^2)$$

$$= \frac{1}{2} m \left[ (\dot{x} + \dot{r} \cos \theta - r \sin \theta \dot{\theta})^2 + (-\dot{r} \sin \theta - r \cos \theta \dot{\theta})^2 \right]$$

$$= \frac{1}{2} m \left[ \dot{x}^2 + \dot{r}^2 \cos^2 \theta + r^2 \sin^2 \theta \dot{\theta}^2 + 2\dot{x}\dot{r} \cos \theta - 2\dot{x}\dot{\theta} r \sin \theta - 2\dot{r}\dot{\theta} \cos \theta \sin \theta + \dot{r}^2 \sin^2 \theta + r^2 \cos^2 \theta \dot{\theta}^2 + 2\dot{r}\dot{\theta} \sin \theta \cos \theta \right]$$

$$T = T_m + T_M$$

$$\mathcal{L} = T - U = \frac{1}{2} (m+M) \dot{x}^2 + \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\theta}^2 + 2\dot{x}\dot{r} \cos \theta - 2\dot{x}\dot{\theta} r \sin \theta) + mgr \sin \theta$$

Constraint  $r - R = 0$ .

To find motion, use  $r = R, \dot{r} = \ddot{r} = 0$ , find Lagrangian eqns for  $\ddot{x}, \ddot{\theta}$ .

Result for  $\ddot{x}$   $\ddot{x} = \frac{mR}{M+m} (\ddot{\theta} \sin \theta + \dot{\theta}^2 \cos \theta)$

(5)

To find constraint force, use an undetermined multiplier,

$$\frac{\partial \mathcal{L}}{\partial q} - \frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \dot{q}} \right) + \lambda \frac{\partial f}{\partial q} = 0$$

$\uparrow$                        $\leftarrow$                        $\leftarrow$   
 Generalized               $\frac{dP}{dt}$                       Constraint  
 force    force

Here, we are interested in the force that keeps  $r=R$ .

$$\frac{\partial \mathcal{L}}{\partial r} - \frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \dot{r}} \right) + \lambda \frac{\partial f}{\partial r} = 0 \quad f = r - R = 0$$

$$\frac{\partial f}{\partial r} = 1$$

$$\frac{1}{2} m \left[ \dot{r}^2 - \dot{x}^2 \sin^2 \theta \right] + mg \sin \theta$$

$$- \frac{d}{dt} \left[ \frac{1}{2} m (\dot{r}^2 + \dot{x}^2 \cos^2 \theta) \right] + \lambda = 0$$

$$m \left( \ddot{r} + \ddot{x} \cos^2 \theta - \dot{x}^2 \sin \theta \dot{\theta} \right)$$

cancel

Now (after taking derivatives), we can impose  $r=R = \text{constraint}$

$$\dot{r} = 0$$

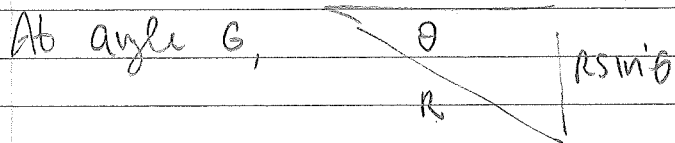
$$m r \dot{\theta}^2 + mg \sin \theta - m \ddot{x} \cos^2 \theta + \lambda = 0$$

You can use Lagrange eqn for  $\ddot{x}$  to get in terms of  $\theta$ .

$$\text{Result: } \lambda = \frac{a-1}{1-a \sin^2 \theta} (R \dot{\theta}^2 + g \sin \theta) \quad a = \frac{m}{M+m}$$

(6)

It's actually easier to use conservation of energy to solve for the motion (than to solve coupled Lagrange eqns.)



$$\star T = mgR \sin \theta = \frac{1}{2} (m+M) \dot{x}^2 + \frac{m}{2} (R^2 \dot{\theta}^2 + 2\dot{x}R\dot{\theta} \sin \theta)$$

$R$  released from top.

$$\text{Now since } \ddot{x} = \frac{mR}{M+m} \ddot{\theta} \sin \theta + \dot{\theta}^2 \cos \theta$$

$$\frac{d}{dt}(\dot{x}) = \frac{mR}{M+m} \left[ \frac{d}{dt}(\dot{\theta}) \cdot \sin \theta + \dot{\theta}^2 \cos \theta \right]$$
$$= \frac{d}{dt}(\dot{\theta} \sin \theta)$$

thus

$$\dot{x} = \frac{mR}{M+m} \dot{\theta} \sin \theta$$

By the way, this is conservation of x momentum!

Now equation  $\star$  can be solved for  $\dot{\theta}^2$  in terms of  $\theta$  and pl'd into eqn. for  $\lambda$ .