

Homework 5 Solutions

1. (a) The dispersion relation is $\omega = \omega_m \left| \sin \frac{1}{2} \mathbf{K} \mathbf{a} \right|$. We solve this for \mathbf{K} to obtain $\mathbf{K} = (2/a) \sin^{-1}(\omega / \omega_m)$, whence $d\mathbf{K}/d\omega = (2/a)(\omega_m^2 - \omega^2)^{-1/2}$ and, from (15), $D(\omega) = (2L/\pi a)(\omega_m^2 - \omega^2)^{-1/2}$. This is singular at $\omega = \omega_m$. (b) The volume of a sphere of radius \mathbf{K} in Fourier space is $\Omega = 4\pi \mathbf{K}^3 / 3 = (4\pi/3)[(\omega_0 - \omega)/A]^{3/2}$, and the density of orbitals near ω_0 is $D(\omega) = (L/2\pi)^3 |d\Omega/d\omega| = (L/2\pi)^3 (2\pi/A^{3/2})(\omega_0 - \omega)^{1/2}$, provided $\omega < \omega_0$. It is apparent that $D(\omega)$ vanishes for ω above the minimum ω_0 .

2. The potential energy associated with the dilation is $\frac{1}{2} B(\Delta V/V)^2 a^3 \approx \frac{1}{2} k_B T$. This is $\frac{1}{2} k_B T$ and not $\frac{3}{2} k_B T$, because the other degrees of freedom are to be associated with shear distortions of the lattice cell. Thus $\langle (\Delta V)^2 \rangle = 1.5 \times 10^{-47}$; $(\Delta V)_{\text{rms}} = 4.7 \times 10^{-24} \text{ cm}^3$; and $(\Delta V)_{\text{rms}}/V = 0.125$. Now $3\Delta a/a \approx \Delta V/V$, whence $(\Delta a)_{\text{rms}}/a = 0.04$.

4. (a) The motion is constrained to each layer and is therefore essentially two-dimensional. Consider one plane of area A . There is one allowed value of \mathbf{K} per area $(2\pi/L)^2$ in \mathbf{K} space, or $(L/2\pi)^2 = A/4\pi^2$ allowed values of \mathbf{K} per unit area of \mathbf{K} space. The total number of modes with wavevector less than \mathbf{K} is, with $\omega = v\mathbf{K}$,

$$N = (A/4\pi^2)(\pi \mathbf{K}^2) = A\omega^2 / 4\pi v^2.$$

The density of modes of each polarization type is $D(\omega) = dN/d\omega = A\omega/2\pi v^2$. The thermal average phonon energy for the two polarization types is, for each layer,

$$U = 2 \int_0^{\omega_D} D(\omega) n(\omega, \tau) \hbar \omega d\omega = 2 \int_0^{\omega_D} \frac{A\omega}{2\pi v^2} \frac{\hbar \omega}{\exp(\hbar \omega / \tau) - 1} d\omega,$$

where ω_D is defined by $N = \int_D^{\omega_D} D(\omega) d\omega$. In the regime $\hbar \omega_D \gg \tau$, we have

$$U \cong \frac{2A\tau^3}{2\pi v^2 \hbar^2} \int_0^\infty \frac{x^2}{e^x - 1} dx.$$

Thus the heat capacity $C = k_B \partial U / \partial \tau \propto T^2$.

- (b) If the layers are weakly bound together, the system behaves as a linear structure with each plane as a vibrating unit. By induction from the results for 2 and 3 dimensions, we expect $C \propto T$. But this only holds at extremely low temperatures such that $\tau \ll \hbar \omega_D \approx \hbar v N_{\text{layer}} / L$, where N_{layer}/L is the number of layers per unit length.