Homework 6 Solutions

1. The energy eigenvalues are $\varepsilon_k = \frac{\hbar^2}{2m} k^2$. The mean value over the volume of a sphere in k space is

$$<\epsilon> = \frac{\text{$\not h$}^2}{2m} \frac{\int k^2 dk \cdot k^2}{\int k^2 dk} = \frac{3}{5} \cdot \frac{\text{$\not h$}^2}{2m} k_F^{\ 2} = \frac{3}{5} \epsilon_F.$$

$$U_0 = N \cdot \frac{3}{5} \varepsilon_F.$$

The total energy of N electrons is . $U_0 = N \cdot \frac{3}{5} \epsilon_F. \label{eq:U0}$

2a. In general $p = -\partial U/\partial V$ at constant entropy. At absolute zero all processes are at constant entropy (the

$$\label{eq:third_law_equation} \text{Third Law), so that } p = -dU_0/dV, \quad \text{where} \quad U_0 = \frac{3}{5} \, N \epsilon_F \\ \quad = \frac{3}{5} \, N \frac{\rlap/ h^2}{2m} \bigg(\frac{3\pi^2 \, N}{V} \bigg)^{2/3}, \quad \text{whence} \\ \quad = \frac{3}{5} \, N \epsilon_F \\ \quad = \frac{3}{5} \, N \frac{\rlap/ h^2}{2m} \bigg(\frac{3\pi^2 \, N}{V} \bigg)^{2/3}, \quad \text{whence} \\ \quad = \frac{3}{5} \, N \epsilon_F \\ \quad = \frac{3}{5} \,$$

$$p = \frac{2}{3} \cdot \frac{U_0}{V}$$
. (b) Bulk modulus

$$B = -V \frac{dp}{dV} = V \left(-\frac{2}{3} \frac{U_0}{V^2} + \frac{2}{3V} \frac{dU_0}{dV} \right) = \frac{2}{3} \cdot \frac{U_0}{V} + \left(\frac{2}{3} \right)^2 \frac{U_0}{V} = \frac{10}{9} \frac{U_0}{V}.$$

(c) For Li,

$$\frac{U_0}{V} = \frac{3}{5} (4.7 \times 10^{22} \text{ cm}^{-3}) (4.7 \text{ eV}) (1.6 \times 10^{-12} \text{ erg/eV})$$
$$= 2.1 \times 10^{11} \text{ erg cm}^{-3} = 2.1 \times 10^{11} \text{ dyne cm}^{-2},$$

whence $B = 2.3 \times 10^{11}$ dyne cm⁻². By experiment (Table 3.3), $B = 1.2 \times 10^{11}$ dyne cm⁻².

3. The number of electrons is, per unit volume, $n = \int_0^\infty d\epsilon \ D(\epsilon) \cdot \frac{1}{e^{(\epsilon - \mu)/\tau} + 1}$, where $D(\epsilon)$ is the density of orbitals. In two dimensions

$$\begin{split} n &= \frac{m}{\pi h^2} \int_0^\infty d\epsilon \frac{1}{e^{(\epsilon - \mu)/\tau} + 1} \\ &= \frac{m}{\pi h^2} (\mu + \tau \log (1 + e^{-\mu/\tau})), \end{split}$$

Use wolfram or equiv to look up integral.

4a. In the sun there are $\frac{2 \times 10^{33}}{1.7 \times 10^{-24}} \approx 10^{57}$ nucleons, and roughly an equal number of electrons. In a white dwarf star of volume

$$\frac{4\pi}{3}(2\times10^9)^3 \approx 3\times10^{28} \text{ cm}^3$$

 $\frac{4\pi}{3}(2\times10^9)^3\approx 3\times10^{28}~cm^3$ the electron concentration is $\approx\frac{10^{57}}{3\times10^{28}}\approx 3\times10^{28}~cm^{-3}$. Thus

$$\epsilon_{\rm F} = \frac{\text{M}^2}{2m} (3\pi^2 n)^{2/3} \approx \frac{1}{2} 10^{-27} \cdot 10^{20} \approx \frac{1}{2} 10^{-7} \text{ ergs, or } \approx 3.10^4 \text{ eV. (b) The value of } k_{\rm F} \text{ is not}$$

affected by relativity and is $\approx n^{1/3}$, where n is the electron concentration. Thus $\epsilon_F \simeq h c k_F \simeq h c^3 \sqrt{n}$. (c) A change of radius to $10 \text{ km} = 10^6 \text{ cm}$ makes the volume $\approx 4 \times 10^{18} \text{ cm}^3$ and the concentration $\approx 3 \times 10^{38} \text{ cm}^{-3}$. Thus $\varepsilon_{\rm F} \approx 10^{-27} (3.10^{10}) (10^{13}) \approx 2.10^{-4} \ erg \approx 10^8 \ eV$. (The energy is relativistic.)